

ORIGINAL ARTICLE

On space-like class $\mathcal{A}$ surfaces in Robertson–Walker spacetimesBurcu Bektaş Demirci¹  | Nurettin Cenk Turgay² | Rüya Yeğin Şen³¹Department of Software Engineering, Faculty of Engineering, Fatih Sultan Mehmet Vakıf University, İstanbul, Turkey²Department of Mathematics, Faculty of Science and Letters, İstanbul Technical University, İstanbul, Turkey³Department of Mathematics, Faculty of Engineering and Natural Sciences, İstanbul Medeniyet University, İstanbul, Turkey

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Abstract

In this paper, we consider space-like surfaces in Robertson–Walker spacetimes $L_1^4(f, c)$ with the comoving observer field $\frac{\partial}{\partial t}$. We study some problems related to such surfaces satisfying the geometric conditions imposed on the tangential and normal parts of the unit vector field $\frac{\partial}{\partial t}$, as naturally defined. First, we investigate space-like surfaces in $L_1^4(f, c)$ satisfying that the tangent component of $\frac{\partial}{\partial t}$ is an eigenvector of all shape operators, called class $\mathcal{A}$ surfaces. Then, we get a classification theorem for space-like class $\mathcal{A}$ surfaces in $L_1^4(f, 0)$. Also, we examine minimal space-like class $\mathcal{A}$ surfaces in $L_1^4(f, 0)$. Finally, we give the parameterizations of space-like surfaces in $L_1^4(f, 0)$ when the normal part of the unit vector field $\frac{\partial}{\partial t}$ is parallel.

KEYWORDS

class $\mathcal{A}$ surfaces, minimal surfaces, Robertson–Walker spacetimes

1 | INTRODUCTION

In recent years, there has been significant interest among geometers in studying submanifolds of product spaces resulting in numerous findings. Some of them are given by [5, 7, 9–11].

Apart from Cartesian product spaces, another example can be considered as Robertson–Walker spacetimes which are four-dimensional Lorentzian manifolds. The Robertson–Walker spacetimes, denoted by $L_1^4(f, c)$, are defined as Cartesian products of space forms by a real interval, equipped with a Lorentzian-warped product metric. Thus, the family of Robertson–Walker spacetimes includes the de Sitter, Minkowski, and the anti-de Sitter spacetimes and also Friedmann’s cosmological models. In physics, Robertson–Walker spacetimes are important due to the fact that they explain homogeneous, isotropic expanding, and contracting universes, see [12, 14].

From geometrical point of view, there have been some studies related to classification of surfaces or hypersurfaces in Robertson–Walker spacetimes, see [1–4, 8]. Especially, Chen and Van der Veken [6] investigated space-like surfaces and Lorentzian surfaces in Robertson–Walker spacetimes having some important geometric properties such as marginally trapped, positive relative nullity, and totally geodesic, etc.

In product spaces, there exists a unit vector field spanning the factor $\mathbb{R}$, denoted by $\frac{\partial}{\partial t}$, and Mendonça and Tojeiro [13] mentioned its decomposition, given as follows.

Given an isometric immersion $\phi : M \rightarrow \mathbb{Q}_c^n \times \mathbb{R}$, a tangent vector field T on M and a normal vector field η along ϕ are defined by

$$\frac{\partial}{\partial t} = \phi_* T + \eta. \quad (1.1)$$

Here, $\mathbb{Q}_c^n$ denotes the space forms. For the Robertson–Walker spacetimes, there still exists the unit vector field, $\frac{\partial}{\partial t}$, which is tangent to first factor and can be decomposed in a similar way to Equation (1.1). Also, this unit vector field is known as a comoving observer field in general relativity [15].

In [16], Tojeiro studied hypersurfaces in product spaces $\mathbb{Q}_c^n \times \mathbb{R}$ for which the tangent component of $\frac{\partial}{\partial t}$ in Equation (1.1) is an eigenvector of all shape operators. Moreover, Mendonça and Tojeiro [13] obtained the characterization of all such submanifolds in the product spaces $\mathbb{Q}_c^n \times \mathbb{R}$. In this context, the following definition for submanifolds of the Cartesian product space, based on [13, 16], can be given.

Definition 1.1. If the vector T in Equation (1.1) is an eigenvector of all shape operators of M , then M is said to be a class $\mathcal{A}$ surface.

In this paper, we study space-like surfaces in Robertson–Walker spacetimes satisfying certain properties in terms of the vectors T and η in the decomposition (1.1). First, we investigate space-like class $\mathcal{A}$ surfaces in $L_1^4(f, c)$ and then, we give a local classification theorem for space-like class $\mathcal{A}$ surfaces in $L_1^4(f, 0)$. Also, we determine space-like class $\mathcal{A}$ surfaces in $L_1^4(f, 0)$ with zero mean curvature vector. Finally, we prove that the vector field η in Equation (1.1) is parallel if and only if the space-like surface in $L_1^4(f, c)$ must be an element of class $\mathcal{A}$ surfaces. Using this, we obtain the parameterizations of space-like surfaces in $L_1^4(f, 0)$ with a parallel vector field η .

2 | PRELIMINARIES

Let $\mathbb{Q}_c^{n-1}$ denote the $n - 1$ dimensional Riemannian space-form with the constant sectional curvature c , that is,

$$\mathbb{Q}_c^{n-1} = \begin{cases} \mathbb{S}^{n-1} & \text{if } c = 1, \\ \mathbb{E}^{n-1} & \text{if } c = 0, \\ \mathbb{H}^{n-1} & \text{if } c = -1 \end{cases}$$

and g_c stand for its metric tensor.

If I is an open interval and $f : I \rightarrow \mathbb{R}$ is a smooth, non-vanishing function, then the Robertson–Walker spacetime $L_1^4(f, c)$ is defined as the Lorentzian warped product $I_1^1 \times_f \mathbb{Q}_c^3$ whose metric tensor $\tilde{g}$ is

$$\tilde{g} = -dt^2 + f^2(t)g_c.$$

Let $\Pi_1 : I \times \mathbb{Q}_c^3 \rightarrow I$ and $\Pi_2 : I \times \mathbb{Q}_c^3 \rightarrow \mathbb{Q}_c^3$ denote the canonical projections. For a given vector field X in $L_1^4(f, c)$, we define a function X_0 and a vector field $\bar{X}$ by the decomposition

$$X = X_0 \frac{\partial}{\partial t} + \bar{X}.$$

Note that we have $X_0 = -\tilde{g}\left(\frac{\partial}{\partial t}, X\right)$ and $\Pi_1^*(\bar{X}) = 0$.

First, we would like to express the Levi-Civita connection of $L_1^4(f, c)$. Note that the following lemma can be directly obtained from [6, Lemma 2.1].

Lemma 2.1. The Levi-Civita connection $\tilde{\nabla}$ of $L_1^4(f, c)$ is

$$\tilde{\nabla}_X Y = \nabla_X^0 Y + (\ln f)'(\tilde{g}(\bar{X}, \bar{Y}))\partial_t + X_0 \bar{Y} + Y_0 \bar{X} \quad (2.1)$$

whenever X and Y are tangent to $L_1^4(f, c)$, where ∇^0 denotes the Levi-Civita connection of the Cartesian product space $L_1^4(1, c) = I \times \mathbb{Q}_c^3$.

2.1 | Space-like surfaces in $L_1^4(f, c)$

First, we give the definition of space-like surfaces in $L_1^4(f, c)$.

Let $\phi : \Omega \rightarrow L_1^4(f, c)$ be a local parameterization of a surface M . Then, M is called a space-like surface if all tangent planes of M are space-like, meaning the metric on these planes is positive definite.

Consider an oriented space-like surface M in $L_1^4(f, c)$ with the Levi-Civita connection ∇ and metric tensor g . Through the misuse of terminology, we shall denote the induced connection on $L_1^4(f, c)$ by $\tilde{\nabla}$. Then, the Gauss and Weingarten formulas

$$\tilde{\nabla}_X Y = \nabla_X Y + h(X, Y), \quad (2.2)$$

$$\tilde{\nabla}_X \xi = -A_\xi(X) + \nabla_X^\perp \xi \quad (2.3)$$

define the second fundamental form h , the shape operator A and the normal connection $\nabla^\perp$ of M , where X, Y are tangent to M and ξ is normal to M . Note that A and h are related by

$$\tilde{g}(h(X, Y), \xi) = g(A_\xi X, Y). \quad (2.4)$$

Remark 2.2. Say $\mathcal{T} = \Pi_1 \circ \phi$. If $\text{grad } \mathcal{T} = 0$ on M , then we have $(M, g) \subset \{t_0\} \times_{f(t_0)} \mathbb{Q}_c^3$. Throughout this work, we are going to exclude this case and assume the existence of $p \in M$ at which $\text{grad } \mathcal{T} \neq 0$.

Since M is a space-like surface, by considering the decomposition (1.1) one may define a function θ and a positively oriented *global* orthonormal frame field $\{e_1, e_2; e_3, e_4\}$ on M by

$$\left. \frac{\partial}{\partial t} \right|_M = \sinh \theta e_1 + \cosh \theta e_3. \quad (2.5)$$

where $g(e_1, e_1) = g(e_2, e_2) = 1$ and $\tilde{g}(e_3, e_3) = -\tilde{g}(e_4, e_4) = -1$. Throughout this paper, we consider such an orthonormal frame field $\{e_1, e_2; e_3, e_4\}$ on M .

Note that we are going to use the notation

$$h_{ij}^\alpha = \tilde{g}(h(e_i, e_j), e_\alpha) = g(e_i, A_{e_\alpha} e_j)$$

for the coefficients of the second fundamental form and ω_{12} will stand for the connection form defined by

$$\omega_{12}(X) = g(\nabla_X e_1, e_2) = -g(\nabla_X e_2, e_1).$$

Then, the mean curvature vector of M is defined by

$$H = \frac{1}{2} \text{tr } h = -\frac{h_{11}^3 + h_{22}^3}{2} e_3 + \frac{h_{11}^4 + h_{22}^4}{2} e_4$$

and M is said to be minimal if H vanishes on M . Also, we are going to use the following lemma:

Lemma 2.3. Let θ and e_1 be as defined above. Then,

$$f'|_M = -\frac{1}{\sinh \theta} e_1(f|_M). \quad (2.6)$$

Proof. Let $p = (t_0, \tilde{p}) \in M$ and consider an integral curve $\alpha = (\alpha_0, \tilde{\alpha})$ of e_1 starting from p . Then, we have

$$e_1(f|_M)_p = \left. \frac{d}{du} \right|_{u=0} (f \circ \alpha)(u) = \left. \frac{d}{du} \right|_{u=0} f(\alpha_0(u)) = \alpha'_0(0) f'(\alpha_0(0))$$

Since $\alpha(0) = p$ and $\alpha'(0) = (e_1)_p$, the last equation implies

$$e_1(f|_M)_p = -\sinh(\theta(p))f'(t_0)$$

which yields equation (2.6). $\square$

During this work, the manifolds that we are dealing with are smooth and connected unless otherwise is stated.

3 | BASIC FACTS FOR SPACE-LIKE SURFACES IN ROBERTSON-WALKER SPACETIMES $L_1^4(f, c)$

In this section, we give some basic facts about space-like surfaces in $L_1^4(f, c)$.

Lemma 3.1. *Let M be a space-like surface in $L_1^4(f, c)$. Then, we have the followings:*

$$\nabla_{e_1} e_1 = h_{12}^3 \coth \theta e_2, \quad (3.1a)$$

$$\nabla_{e_2} e_1 = (\ln f)' \operatorname{csch} \theta + h_{22}^3 \coth \theta e_2, \quad (3.1b)$$

$$\nabla_{e_1}^\perp e_3 = -h_{11}^4 \tanh \theta e_4, \quad (3.1c)$$

$$\nabla_{e_2}^\perp e_3 = -h_{12}^4 \tanh \theta e_4 \quad (3.1d)$$

$$e_1(\theta) = (\ln f)' \cosh \theta + h_{11}^3, \quad (3.1e)$$

$$e_2(\theta) = h_{12}^3. \quad (3.1f)$$

Proof. Assume that M is a space-like surface in $L_1^4(f, c)$. From equation (2.5), we get $(e_2)_0 = \tilde{g}\left(e_2, \frac{\partial}{\partial t}\right) = 0$ and $(e_1)_0 = -\tilde{g}\left(\frac{\partial}{\partial t}, e_1\right) = -\sinh \theta$. Then, we have $e_2 = \bar{e}_2$ and $e_1 = -\sinh \theta \frac{\partial}{\partial t} + \bar{e}_1$. Considering equation (2.1) in Lemma 2.1 with these equations, one can obtain

$$\tilde{\nabla}_{e_1} \frac{\partial}{\partial t} = (\ln f)' \left(\cosh^2 \theta e_1 + \sinh \theta \cosh \theta e_3 \right), \quad (3.2)$$

$$\tilde{\nabla}_{e_2} \frac{\partial}{\partial t} = (\ln f)' e_2. \quad (3.3)$$

On the other hand, Equation (2.5) implies that

$$\tilde{\nabla}_X \frac{\partial}{\partial t} = X(\theta) \cosh \theta e_1 + \sinh \theta \tilde{\nabla}_X e_1 + X(\theta) \sinh \theta e_3 + \cosh \theta \tilde{\nabla}_X e_3$$

from which we obtain

$$\begin{aligned} \left(\tilde{\nabla}_X \frac{\partial}{\partial t} \right)^T &= X(\theta) \cosh \theta e_1 + \sinh \theta \nabla_X e_1 - \cosh \theta A_{e_3} X, \\ \left(\tilde{\nabla}_X \frac{\partial}{\partial t} \right)^\perp &= X(\theta) \sinh \theta e_3 + \sinh \theta h(X, e_1) + \cosh \theta \nabla_X^\perp e_3. \end{aligned} \quad (3.4)$$

Comparing Equation (3.4) for $X = e_1$ and Equation (3.2), we get equations (3.1a), (3.1c), and (3.1e). Similarly, equation (3.4) for $X = e_2$ and Equation (3.3) give Equations (3.1b), (3.1d), and (3.1f). $\square$

From Lemma 3.1, we give the following proposition.

Proposition 3.2. *Let M be a space-like surface in $L_1^4(f, c)$. Then, the following statements are equivalent to each other:*

- (i) M is space-like class $\mathcal{A}$ surface in $L_1^4(f, c)$.
- (ii) For the vector fields e_1, e_2 tangent to M , we have $h(e_1, e_2) = 0$.
- (iii) For the vector fields e_2, e_4 , we have $e_2(\theta) = 0$ and $\nabla_{e_2}^\perp e_4 = 0$.

Proof. From Definition 1.1 and Equation (2.4), it can be easily seen that statements (i) and (ii) are equivalent to each other. Statements (ii) and (iii) are direct consequences of Equations (3.1d) and (3.1f). $\square$

Lemma 3.3. *Let M be a space-like class $\mathcal{A}$ surface in $L_1^4(f, c)$ and $p \in M$. Then, there exists a local coordinate system (u, v) defined on a neighborhood $\mathcal{N}_p$ of p which can be parameterized by*

$$\phi(u, v) = (u, \tilde{\phi}(u, v)) \quad (3.5)$$

for an immersion $\tilde{\phi} : \Omega \subset \mathbb{R}^2 \rightarrow \mathbb{Q}_c^3$ satisfying

$$g_c(\tilde{\phi}_u, \tilde{\phi}_v) = 0 \quad \text{and} \quad \partial_v(g_c(\tilde{\phi}_u, \tilde{\phi}_u)) = 0. \quad (3.6)$$

Proof. Assume that M is a space-like class $\mathcal{A}$ surface in $L_1^4(f, c)$. Then, Equation (3.1a) becomes $\nabla_{e_1} e_1 = 0$. Define $X = -\frac{1}{\sinh \theta} e_1$ and $Y = \gamma e_2$. Then, we have

$$[X, Y] = \operatorname{csch} \theta (-e_1(\gamma) + \gamma((\ln f)' \operatorname{csch} \theta + h_{22}^3 \coth \theta)) e_2. \quad (3.7)$$

Thus, we choose a smooth function γ satisfying $e_1(\gamma) = \gamma \omega$ which makes $[X, Y] = 0$. Hence, there exists a neighborhood $\mathcal{N}_p$ of p on which a local coordinate system (u, v) is defined such that

$$X = -\frac{1}{\sinh \theta} e_1 = \partial_u, \quad Y = \gamma e_2 = \partial_v. \quad (3.8)$$

Now, we consider a parameterization of M in $L_1^4(f, c)$

$$\phi(u, v) = (\mathcal{T}(u, v), \tilde{\phi}(u, v)), \quad (3.9)$$

where $\operatorname{grad} \mathcal{T} \neq 0$ and $\tilde{\phi} : \Omega \subset \mathbb{R}^2 \rightarrow \mathbb{Q}_c^3$ is an immersion. From the equations $\tilde{g}\left(e_1, \frac{\partial}{\partial t}\right) = \sinh \theta$ and $\tilde{g}\left(e_2, \frac{\partial}{\partial t}\right) = 0$, we have

$$\frac{\partial \mathcal{T}}{\partial u} = 1, \quad \frac{\partial \mathcal{T}}{\partial v} = 0$$

on $\mathcal{N}_p$, respectively. Thus, M has a parameterization given by Equation (3.5) with $g_c(\tilde{\phi}_u, \tilde{\phi}_v) = 0$. Moreover, we get

$$\tilde{g}(\phi_u, \phi_u) = -1 + f^2 g_c(\tilde{\phi}_u, \tilde{\phi}_u) = \operatorname{csch}^2 \theta \quad (3.10)$$

Considering $e_2(\theta) = 0$ from Proposition 3.2 and Equation (3.10), we have $\partial_v(g_c(\tilde{\phi}_u, \tilde{\phi}_u)) = 0$. Thus, we get the desired result. $\square$

The proof of the following corollary directly follows from the proof of Lemma 3.3.

Corollary 3.4. *A space-like surface in $L_1^4(f, c)$ satisfies $A_\eta T = \lambda T$ for a smooth function λ and a normal vector η if and only if it can be locally parameterized by Equation (3.5) for an immersion $\tilde{\phi}$ satisfying Equation (3.6).*

4 | CLASS $\mathcal{A}$ SURFACES IN $L_1^4(f, 0)$

In this section, we will give the local classification theorem of space-like class $\mathcal{A}$ surfaces for $c = 0$. Thus, we focus on the surface M given in Lemma 3.3.

4.1 | Local classification theorem

Let $\tilde{M}$ denote the surface in $\mathbb{E}^3$ parameterized by $\tilde{\phi}$ with a unit normal vector field $\tilde{N}$ and $\tilde{E}, \tilde{G}$ stand for the coefficients of the first fundamental form of $\tilde{M}$, i.e., $\tilde{E} = g_0(\tilde{\phi}_u, \tilde{\phi}_u)$ and $\tilde{G} = g_0(\tilde{\phi}_v, \tilde{\phi}_v)$.

First, we construct a geometrical frame field on M as follows.

Lemma 4.1. *Let M be a space-like surface in $L_1^4(f, 0)$ parameterized by Equation (3.5). Then, there exists an orthonormal frame field $\{e_1, e_2, e_3, e_4\}$ on M given by*

$$\begin{aligned} e_1 &= \frac{1}{\sqrt{-1 + f^2 \tilde{E}}} \partial_u, & e_2 &= \frac{1}{f \sqrt{\tilde{G}}} \partial_v, \\ e_3 &= \frac{1}{\sqrt{f^2 \tilde{E}^2 - \tilde{E}}} \left(f \tilde{E}, \frac{1}{f} \tilde{\phi}_u \right), \\ e_4 &= \frac{1}{f} (0, \tilde{N}). \end{aligned} \quad (4.1)$$

Moreover, the functions θ and h_{12}^4 satisfy

$$\theta = -\sinh^{-1} \left(\frac{1}{\sqrt{-1 + f^2 \tilde{E}}} \right), \quad (4.2)$$

$$h_{12}^4 = \frac{f}{\sqrt{EG}} g_0(\tilde{\phi}_{uv}, \tilde{N}). \quad (4.3)$$

Proof. By considering the parameterization of the immersion ϕ in Equation (3.5) and the first condition in Equation (3.6), one can observe that the vector fields e_1, e_2, e_3, e_4 are as given in (4.1). From Lemma 2.3 and Equation (4.1) we have $f'(t) = f'(u)$ whenever $(t, \bar{p}) = \phi(u, v) \in M$. Therefore, by a direct computation using Equations (2.1) and (4.1) we get desired results. $\square$

Now, we consider the following examples of space-like class $\mathcal{A}$ surfaces in $L_1^4(f, 0)$.

Example 4.2. For some smooth functions x_1, x_2 , consider the following space-like surface in $L_1^4(f, 0)$:

$$\phi(u, v) = (u, x_1(u), x_2(u), v) \quad (4.4)$$

with $-1 + f^2(x_1'^2(u) + x_2'^2(u)) > 0$. Say $V(u) = \sqrt{x_1'^2(u) + x_2'^2(u)}$. Then, Equation (4.1) turns into

$$\begin{aligned} e_1 &= \frac{1}{\sqrt{-1 + f^2 V^2}} \partial_u, & e_2 &= \frac{1}{f} \partial_v, \\ e_3 &= \frac{1}{f V \sqrt{f^2 V^2 - 1}} (f^2 V^2, x_1', x_2', 0), \\ e_4 &= \frac{1}{f V} (0, -x_2', x_1', 0) \end{aligned} \quad (4.5)$$

and $\sinh \theta = -\frac{1}{\sqrt{-1 + f^2 V^2}}$. Thus, it can be seen that $e_2(\theta) = 0$ and $\nabla_{e_2}^\perp e_4 = 0$. From Proposition 3.2, we get that M is a space-like class $\mathcal{A}$ surface in $L_1^4(f, 0)$. For later use, we also give the nonzero components of the second fundamental form h of M as

$$\begin{aligned} h_{11}^3 &= \frac{fV' - V(f^2 V^2 - 2)f'}{(f^2 V^2 - 1)^{3/2}}, & h_{22}^3 &= -\frac{Vf'}{\sqrt{f^2 V^2 - 1}}, \\ h_{11}^4 &= \frac{f(x_2' x_1'' - x_1' x_2'')}{V - f^2 V^3}, & h_{22}^4 &= 0. \end{aligned} \quad (4.6)$$

Example 4.3. Let $\alpha : I_v \rightarrow \mathbb{S}^2$ be an arc-length parameterized curve with unit normal n and curvature κ . Consider the following space-like surface in $L_1^4(f, 0)$ given by

$$\phi(u, v) = (u, \phi_1(v)\alpha(v) + \phi_2(u, v)\alpha'(v) + \phi_3(u, v)n(v)) \quad (4.7)$$

where $\phi_1(v)$, $\phi_2(u, v)$, and $\phi_3(u, v)$ are smooth functions defined by

$$\begin{aligned} \phi_2(u, v) &= \int_{u_0}^u R(\xi) \sin(\tau(\xi, v)) d\xi + \psi_1(v) \\ \phi_3(u, v) &= \int_{u_0}^u R(\xi) \cos(\tau(\xi, v)) d\xi + \psi_2(v) \end{aligned} \quad (4.8)$$

for some smooth functions R , τ , ψ_1 , and ψ_2 satisfying $\tau_v = \kappa$ and

$$\psi_1' = \kappa\psi_2 - \phi_1, \quad \psi_2' = -\kappa\psi_1. \quad (4.9)$$

Since it is space-like, we have $(-1 + f^2 R^2)(\phi_1' - \phi_2) > 0$. By a direct computation, one can obtain that

$$e_2 = \frac{1}{f(\phi_1' - \phi_2)} \partial_v, \quad e_4 = \frac{1}{f} (0, \cos \tau \alpha' - \sin \tau n), \quad (4.10)$$

$$\sinh \theta = -\frac{1}{\sqrt{-1 + f^2 R^2}}. \quad (4.11)$$

It can be seen that $e_2(\theta) = 0$ and $\nabla_{e_2}^\perp e_4 = 0$. Thus, Proposition (3.2) implies that M is a space-like class $\mathcal{A}$ surface in $L_1^4(f, 0)$.

Moreover, if $\kappa = 0$, then the surface described in Example 4.3 turns into the surface of revolution given by the following example.

Example 4.4. Consider the following space-like surface parameterized by

$$\phi(u, v) = (u, \zeta_1(u) \cos v, \zeta_1(u) \sin v, \zeta_2(u)) \quad (4.12)$$

for some smooth functions $\zeta_1 > 0$ and ζ_2 . Let define $V(u) = \sqrt{\zeta_1'^2(u) + \zeta_2'^2(u)}$. Then, Equation (4.1) turns into

$$\begin{aligned} e_1 &= \frac{1}{\sqrt{-1 + f^2 V^2}} \partial_u, \quad e_2 = \frac{1}{f \zeta_1} \partial_v, \\ e_3 &= \frac{1}{f V \sqrt{-1 + f^2 V^2}} (f^2 V^2, \zeta_1' \cos v, \zeta_1' \sin v, \zeta_2'), \\ e_4 &= \frac{1}{f V} (0, \zeta_2' \cos v, \zeta_2' \sin v, -\zeta_1'), \end{aligned}$$

By a direct computation, we obtain the nonzero coefficients of the second fundamental form h of M as

$$\begin{aligned} h_{11}^3 &= \frac{V(2 - f^2 V^2)f' + fV'}{(-1 + f^2 V^2)^{3/2}}, \quad h_{22}^3 = \frac{V^2 \sqrt{-1 + f^2 V^2} f' + \zeta_1'}{V - f^2 V^3}, \\ h_{11}^4 &= \frac{f(\zeta_2' \zeta_1'' - \zeta_1' \zeta_2'')}{V(-1 + f^2 V^2)}, \quad h_{22}^4 = -\frac{\zeta_2'}{V \sqrt{-1 + f^2 V^2}}. \end{aligned} \quad (4.13)$$

Now, we give the local classification of space-like class $\mathcal{A}$ surfaces in $L_1^4(f, 0)$.

Theorem 4.5. A space-like surface in $L_1^4(f, 0)$ is a class $\mathcal{A}$ surface if and only if it is locally congruent to one of the following surfaces:

- (i) The cylinder described in Example 4.2,
- (ii) The surface described in Example 4.3.

Proof. Assume that M is a space-like class $\mathcal{A}$ surface in $L_1^4(f, 0)$ and let $p \in M$. Then, Lemma 3.3 implies that there exists local coordinates (u, v) on a neighborhood $\mathcal{N}_p$ of p which can be parameterized by Equation (3.5) immersion $\tilde{\phi}$ satisfying Equation (3.6). Then, from Proposition 3.2, we have $\tilde{g}(h(e_1, e_2), e_4) = h_{12}^4 = 0$. Thus, Equation (4.3) gives $g_0(\tilde{\phi}_{uv}, \tilde{N}) = 0$. Therefore, we express

$$\tilde{\phi}_{uv} = \frac{g_0(\tilde{\phi}_{uv}, \tilde{\phi}_v)}{g_0(\tilde{\phi}_v, \tilde{\phi}_v)} \tilde{\phi}_v = \frac{\tilde{G}_u}{2\tilde{G}} \tilde{\phi}_v$$

from which we get

$$\tilde{\phi}_v = \sqrt{\tilde{G}(u, v)} \alpha(v) \quad (4.14)$$

for an $\mathbb{R}^3$ -valued function α satisfying $g_0(\alpha, \alpha) = 1$. Then, the following two cases occur depending on the function α .

Case (i.) α is a constant vector in $\mathbb{R}^3$. In this case, up to a suitable isometry, we may assume $\alpha = (0, 0, 1)$. Considering Equation (4.14), the case (i) of the theorem can be obtained.

Case (ii.) α is a nonconstant vector in $\mathbb{R}^3$, that is, $\alpha' \neq 0$. In this case, by re-defining v properly, one may assume $g_0(\alpha', \alpha') = 1$. Thus, $\alpha : I_v \rightarrow \mathbb{S}^2$ is an arc-length parameterized spherical curve. Let κ and n be a curvature and a normal vector of α , where we have $\alpha'' = \kappa n - \alpha$ and $n' = -\kappa \alpha'$. Then, the immersion $\tilde{\phi}$ in $\mathbb{E}^3$ can be written as

$$\tilde{\phi}(u, v) = \phi_1(u, v)\alpha(v) + \phi_2(u, v)\alpha'(v) + \phi_3(u, v)n(v) \quad (4.15)$$

for some smooth functions ϕ_1, ϕ_2, ϕ_3 . Since $\tilde{g}(\tilde{\phi}_u, \tilde{\phi}_v) = 0$, Equations (4.14) and (4.15) give $\frac{\partial \phi_1}{\partial u} = 0$ which implies $\phi_1 = \phi_1(v)$. Considering Equation (4.2) and $e_2(\theta) = 0$ together, we have $\tilde{E}(u, v) = R^2(u)$ for a non-vanishing function R . Thus, we get

$$\frac{\partial \phi_2}{\partial u} = R \sin \tau, \quad \frac{\partial \phi_3}{\partial u} = R \cos \tau, \quad (4.16)$$

for a smooth function $\tau = \tau(u, v)$. Integrating Equation (4.16), we get Equations (4.8) for ϕ_2 and ϕ_3 with some functions ψ_1 and ψ_2 . On the other hand, by a direct computation using Equations (4.7), (4.16) and $g_0(\tilde{\phi}_u, \tilde{\phi}_v) = 0$, we obtain

$$\frac{\partial \phi_2}{\partial v} + \phi_1 - \kappa \phi_3 = 0, \quad \frac{\partial \phi_3}{\partial v} + \kappa \phi_2 = 0. \quad (4.17)$$

Taking the derivative of Equations (4.16) and (4.17) with respect to v and u , we get $\frac{\partial \tau}{\partial v} = \kappa$ and Equation (4.8) for some smooth functions ψ_1 and ψ_2 . By using Equations (4.8) and (4.17), we obtain Equation (4.9). Thus, M is congruent to the surface given in case (ii) of the theorem. Hence, the proof of the necessary condition is completed.

The proof of the sufficient condition is obtained from Examples 4.2 and 4.3. □

4.2 | Minimal surfaces

In this subsection, as an application of Theorem 4.5 we study minimal class $\mathcal{A}$ surfaces in $L_1^4(f, 0)$. First, we will focus on the surface given in (i) of Theorem 4.5.

Proposition 4.6. *The class $\mathcal{A}$ space-like surface described in (i) of Theorem 4.5 is minimal if and only if it is congruent to the surface given by*

$$\phi(u, v) = (u, x_1(u), c_1 x_1(u) + c_2, v) \quad (4.18)$$

for constants c_1 , c_2 and a function x_1 satisfying

$$x_1 = \int_{u_0}^u \frac{d\xi}{f(\xi)\sqrt{c_3 f^4(\xi) + (c_1^2 + 1)}} \quad (4.19)$$

where u_0, c_3 are real constants.

Proof. Assume that M is a minimal class $\mathcal{A}$ surface described in (i) of Theorem 4.5, that is, $h_{11}^3 + h_{22}^3 = h_{11}^4 + h_{22}^4 = 0$. From the equations in Equation (4.6), we find $x_2'x_1'' - x_1'x_2'' = 0$ whose solution is given by $x_2(u) = c_1x_1(u) + c_2$ for some constants c_1, c_2 . By a further computation, considering $h_{11}^3 + h_{22}^3 = 0$ and Equation (4.6) we obtain the following differential equation:

$$fx_1'' - 2f^2(1 + c_1^2)f'x_1'^3 + 3f'x_1' = 0.$$

Then, the solution of this equation is given by Equation (4.19). $\square$

Remark 4.7. The surface given by Equation (4.18) lies on a totally geodesic hypersurface of $L_1^4(f, 0)$.

Now, we will examine the surface given in (ii) of Theorem 4.5 with minimality condition.

Lemma 4.8. *The class $\mathcal{A}$ space-like surface described in (ii) of Theorem 4.5 is a minimal surface if and only if $\kappa = 0$.*

Proof. Assume that M is a minimal class $\mathcal{A}$ surface described in (ii) of Theorem 4.5. Let $\bar{M}$ denote the surface in $\mathbb{E}^3$ parameterized by

$$\tilde{\phi}(u, v) = \phi_1(v)\alpha(v) + \phi_2(u, v)\alpha'(v) + \phi_3(u, v)n(v).$$

By a direct computation, we obtain that the principal directions of $\bar{M}$ are $\bar{e}_1 = \frac{1}{R}\tilde{\phi}_*(\partial_u)$ and $\bar{e}_2 = \frac{1}{\phi_1' - \phi_2}\tilde{\phi}_*(\partial_v)$ with the corresponding principal curvatures k_1, k_2 given by

$$k_1 = \frac{\tau_u}{R}, \quad k_2 = \frac{\cos \tau}{\phi_1' - \phi_2}. \quad (4.20)$$

Since $\tau_v = \kappa(v)$, we get $\bar{e}_2(k_1) = 0$ which means $k_1 = k_1(u)$. By a direct computation, we obtain that k_1 and k_2 are related with

$$\bar{e}_1(k_2) = \frac{\sin \tau}{\phi_1' - \phi_2}(k_2 - k_1). \quad (4.21)$$

On the other hand, Equation (4.1) implies

$$e_4 = \frac{1}{f}(0, \cos \tau \alpha' - \sin \tau n). \quad (4.22)$$

Considering Equations (4.20) and (4.22), we obtain

$$h_{11}^4 + h_{22}^4 = \frac{fR^2k_1}{-1 + f^2R^2} + \frac{k_2}{f}. \quad (4.23)$$

Since M is a minimal surface, that is, $h_{11}^4 + h_{22}^4 = 0$, $k_2 = \frac{f^2R^2k_1}{1 - f^2R^2}$ which implies $k_2 = k_2(u)$. Considering this with Equations (4.20) and (4.21), we obtain $\tau_v = \kappa = 0$. $\square$

Hence, we give the following proposition:

Proposition 4.9. *The class $\mathcal{A}$ surface described in (ii) of Theorem 4.5 is minimal if and only if it is congruent to the surface given by Equation (4.12) in Example 4.4 for some smooth functions ζ_1 and ζ_2 satisfying*

$$\begin{aligned} f\zeta_1'' &= f'\zeta_1'(2f^2(\zeta_1'^2 + \zeta_2'^2) - 3) + \sqrt{f^2(\zeta_1'^2 + \zeta_2'^2) - 1}, \\ f\zeta_2'' &= f'\zeta_2'(2f^2(\zeta_1'^2 + \zeta_2'^2) - 3). \end{aligned} \quad (4.24)$$

Proof. Suppose that M is a class $\mathcal{A}$ surface described in (ii) of Theorem 4.5 and it is minimal. Thus, Lemma 4.8 implies that M is congruent to the surface parameterized by Equation (4.12) for some smooth functions $\zeta_1 > 0$ and ζ_2 . By considering Equation (4.13), we obtain the equations in Equation (4.24). $\square$

By combining Proposition 4.6, Remark 4.7, and Proposition 4.9, we obtain the following result:

Theorem 4.10. *Let M be a space-like surface in $L_1^4(f, 0)$ which has no open part lying on a totally geodesic hypersurface of $L_1^4(f, 0)$. Then, M is minimal and class $\mathcal{A}$ surface if and only if it is locally congruent to the surface (4.12) for some smooth functions ζ_1 and ζ_2 satisfying Equation (4.24).*

5 | SPACE-LIKE SURFACES IN $L_1^4(f, 0)$ SATISFYING $\nabla^\perp \eta = 0$

In this section, we consider space-like surfaces in $L_1^4(f, 0)$ having a parallel normal vector field η in Equation (1.1). First, we will give the following lemma to obtain the classification theorem.

Lemma 5.1. *Let M be a space-like surface in $L_1^4(f, c)$. Then, the vector field η in Equation (1.1) is parallel if and only if M is a class $\mathcal{A}$ surface in $L_1^4(f, c)$.*

Proof. Suppose that M is a space-like surface in $L_1^4(f, c)$ and $\nabla_X^\perp \eta = 0$ for any tangent vector field X . From Equation (2.5), we know that $\eta = \cosh \theta e_3$. Thus, we get that θ is a nonzero constant and $\nabla_X^\perp e_3 = 0$. Note that when θ is zero, $\frac{\partial}{\partial t} = e_3$ which means $T = 0$. We omit this case. From Proposition 3.2, it can be seen that M is a class $\mathcal{A}$ surface in $L_1^4(f, c)$. $\square$

By using Lemma 5.1, we will state the following theorem.

Theorem 5.2. *A space-like surface in $L_1^4(f, 0)$ having the parallel vector field η in Equation (1.1) is locally congruent to one of the following surfaces:*

(i) *The cylinder given by*

$$\phi(u, v) = \left(u, c_1 \int_{u_0}^u \frac{d\xi}{f(\xi)}, c_2 \int_{u_0}^u \frac{d\xi}{f(\xi)} + c_3, v \right) \quad (5.1)$$

for some constants u_0 , c_1 , c_2 and c_3 .

(ii) *The surface parameterized by Equation (4.7), where ϕ_1 is a smooth function and the functions ϕ_2, ϕ_3 are given by*

$$\begin{aligned} \phi_2(u, v) &= c \sin \tau(v) \int_{u_0}^u \frac{d\xi}{f(\xi)} + \psi_1(v) \\ \phi_3(u, v) &= c \cos \tau(v) \int_{u_0}^u \frac{d\xi}{f(\xi)} + \psi_2(v) \end{aligned} \quad (5.2)$$

for constants c, u_0 , smooth functions τ, ψ_1 and ψ_2 satisfying Equation (4.9) and $\tau'(v) = \kappa(v)$.

Proof. Suppose that M is a space-like surface in $L_1^4(f, 0)$ with parallel vector field η . Then, Lemma (5.1) implies that M is a class $\mathcal{A}$ surface. Thus, we have two surfaces given in Theorem 4.5. Now, we are going to study these surfaces, separately.

Case (i). M is congruent to the cylinder described in Example 4.2 for some smooth functions x_1, x_2 . In this case, since θ is constant, we have

$$x_1'^2(u) + x_2'^2(u) = \frac{c^2}{f^2(u)} \quad (5.3)$$

for a constant $c > 0$. From Equation (3.1c), we get $h_{11}^4 = 0$. Thus, Equation (4.6) gives $x_2'x_1'' - x_1'x_2' = 0$ which yields $x_2(u) = a_1x_1(u) + a_2$ for some constants a_1, a_2 . By combining this equation with Equation (5.3), we get $x_1' = \frac{c_1}{f}$. Then, Equation (4.4) turns into Equation (5.1). Hence, we have the case (i) of the theorem.

Case (ii). The surface described in Example 4.3. In this case, similar to the case (i), we have

$$R = \frac{c}{f}. \quad (5.4)$$

Using Equation (4.22) and $h_{11}^4 = 0$, we get $\tau_u = 0$. Thus, Equation (4.8) turns into Equation (5.2). Hence, we have case (ii) of the theorem.

The converse can be shown through a straightforward calculation. □

6 | CONCLUSION

In this paper, we study space-like surfaces in Robertson–Walker spacetimes with certain properties T and η vectors in the decomposition (1.1). First, we examine space-like surfaces in $L_1^4(f, c)$ satisfying that T is an eigenvector of all shape operators, called class $\mathcal{A}$ surfaces. Then, we give a local classification theorem for space-like class $\mathcal{A}$ surfaces in $L_1^4(f, 0)$ whose fiber is $\mathbb{R}^3$. As a particular case, we get minimal space-like class $\mathcal{A}$ surfaces in $L_1^4(f, 0)$. Finally, we prove that the vector field η in Equation (1.1) is parallel if and only if the space-like surface in $L_1^4(f, c)$ belongs to the family of class $\mathcal{A}$ surfaces. Thus, we obtain the parameterizations of space-like surfaces in $L_1^4(f, 0)$ with a parallel vector field η .

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CONFLICT OF INTEREST STATEMENT

The authors declare no conflicts of interest.

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